

# Neutrino factories & muon colliders at the energy frontier

**R. B. Palmer**

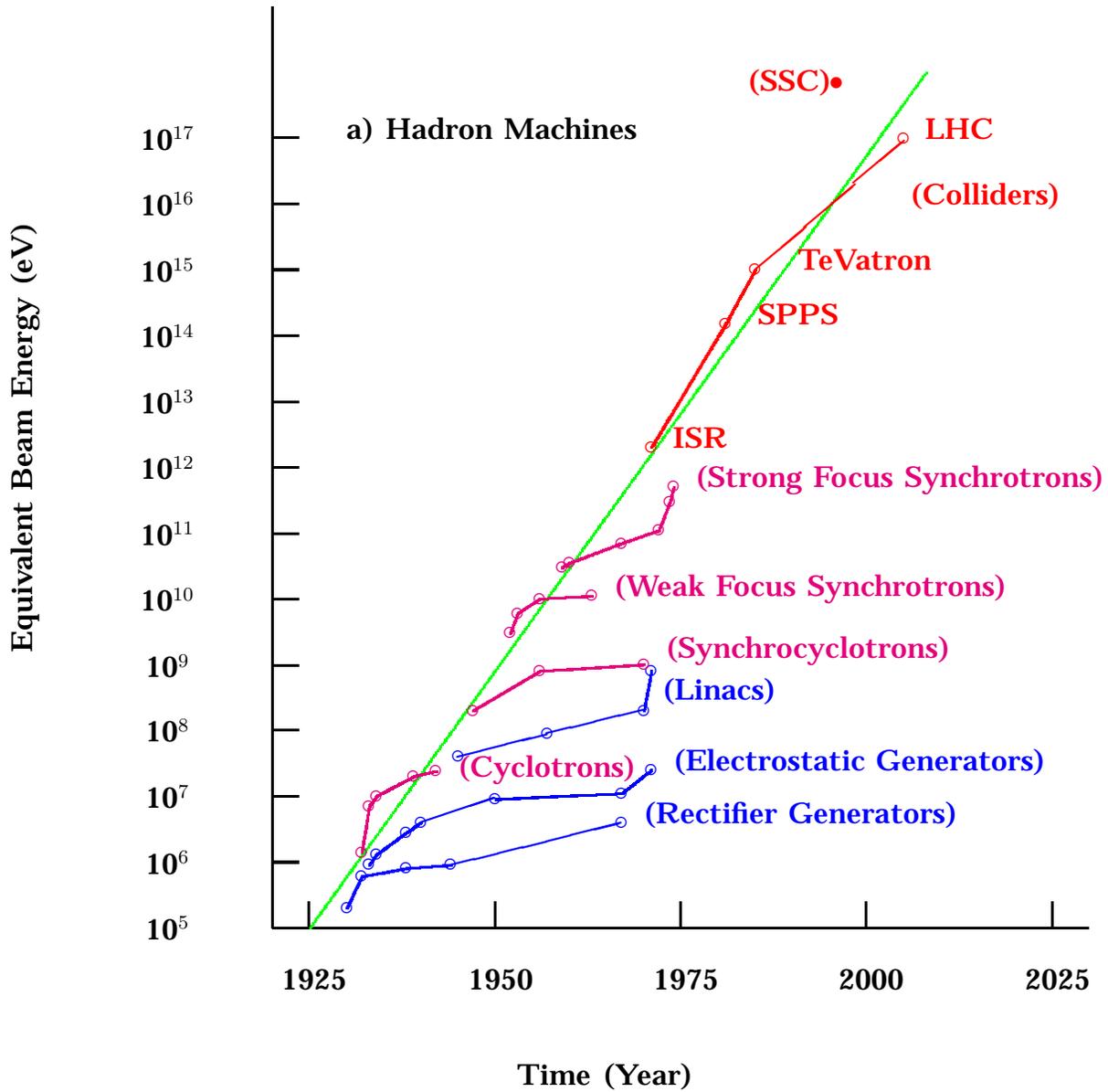
Brookhaven Nat. Lab

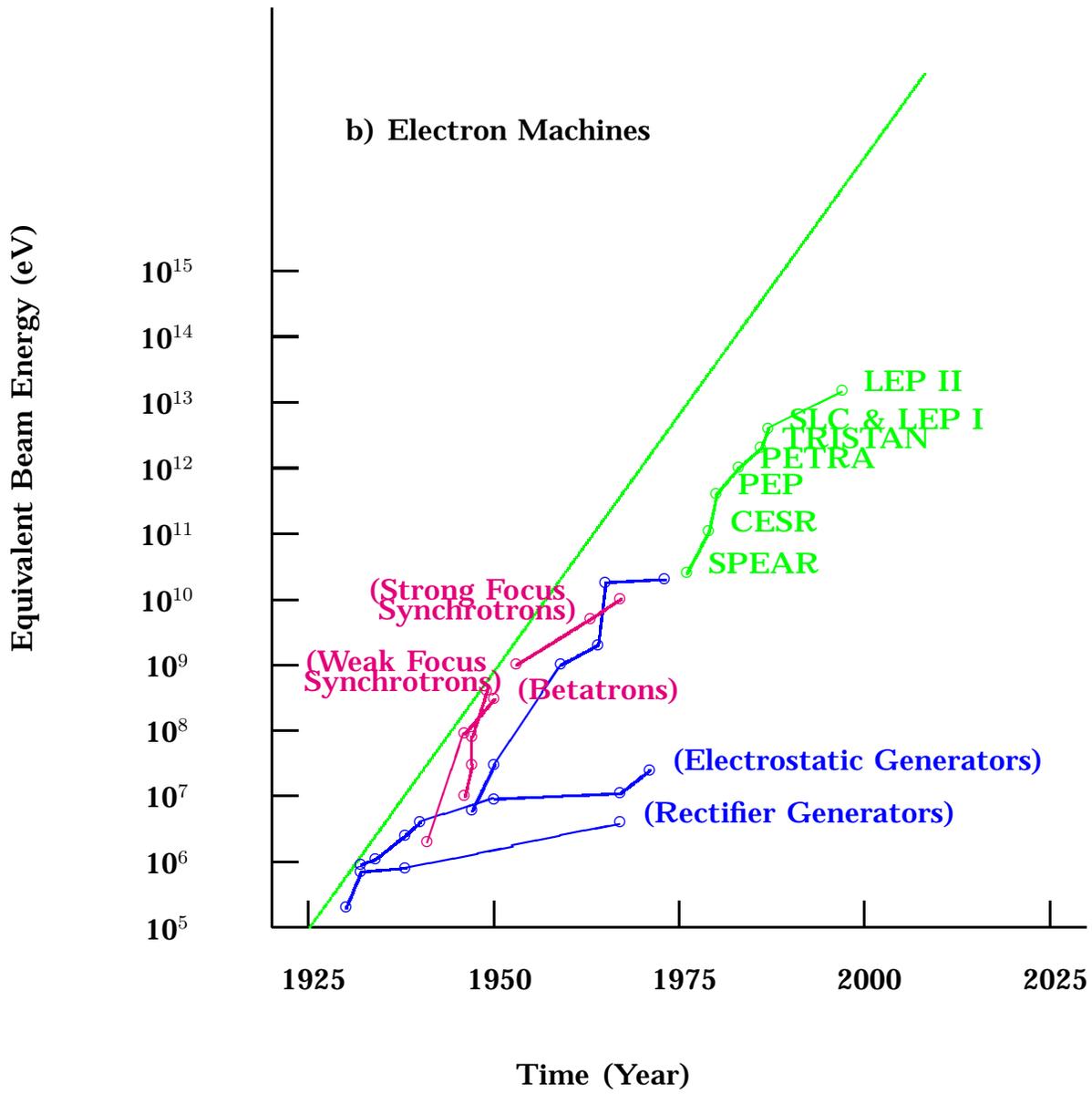
Oxford

November 2001

- **Electron Colliders**
- **Hadron Colliders**
- **Muon Colliders**
- **Neutrino Factories**

# LIVINGSTON PLOTS





## $e^+ e^-$ COLLIDERS

- E advantage over protons (5-10)
- Known initial state
- Low backgrounds & radiation
- Complements protons
- Must be linear for  $E > .2$  TeV

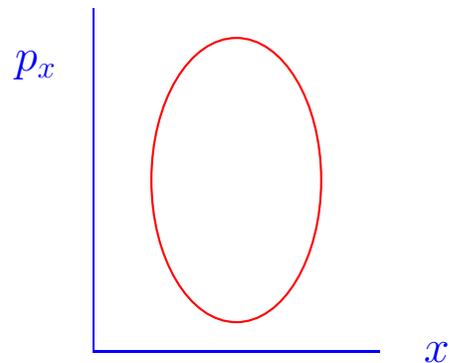
## WHAT IS LUMINOSITY

$$L = \frac{\text{event rate}}{\text{interaction cross section}}$$

## WHAT IS EMITTANCE

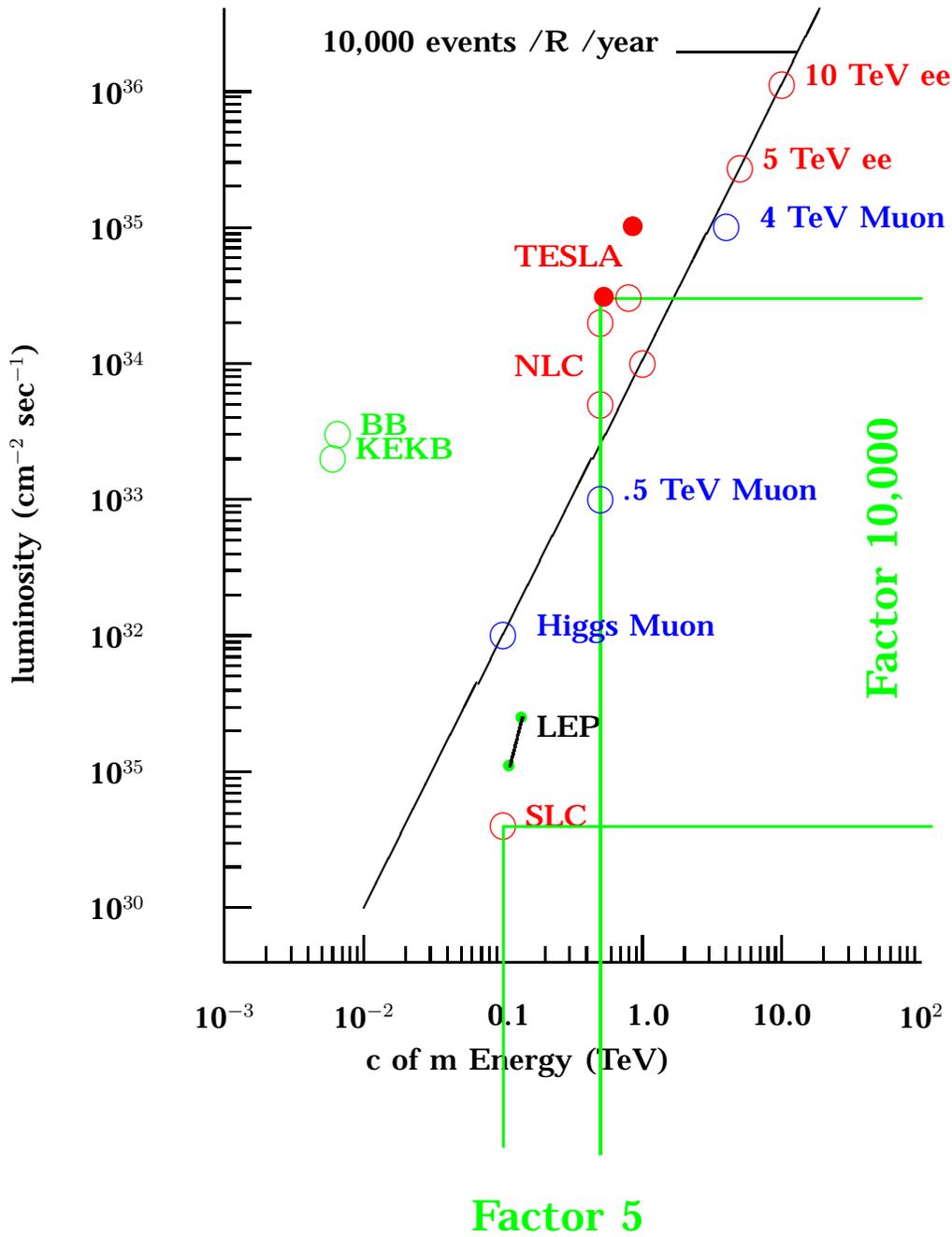
$$\text{normalized emittance} = \frac{\text{Phase Space Area}}{m c}$$

If  $x$  and  $p_x$  both Gaussian and uncorrelated, then area is an upright ellipse



$$\epsilon_{\perp} = \frac{dp_{\perp} dx}{mc} = \sigma_{\theta} \sigma_x (\gamma \beta_v) \quad (\pi \text{ m rad})$$

# LUMINOSITY



## WHY IS LUMINOSITY HARD ?

$$L \propto \frac{P_{ave}}{\gamma} \left( \frac{N}{\sigma_x} \right) \left( \frac{1}{\epsilon_n^{(5/7)}} \right) \propto \gamma^2$$

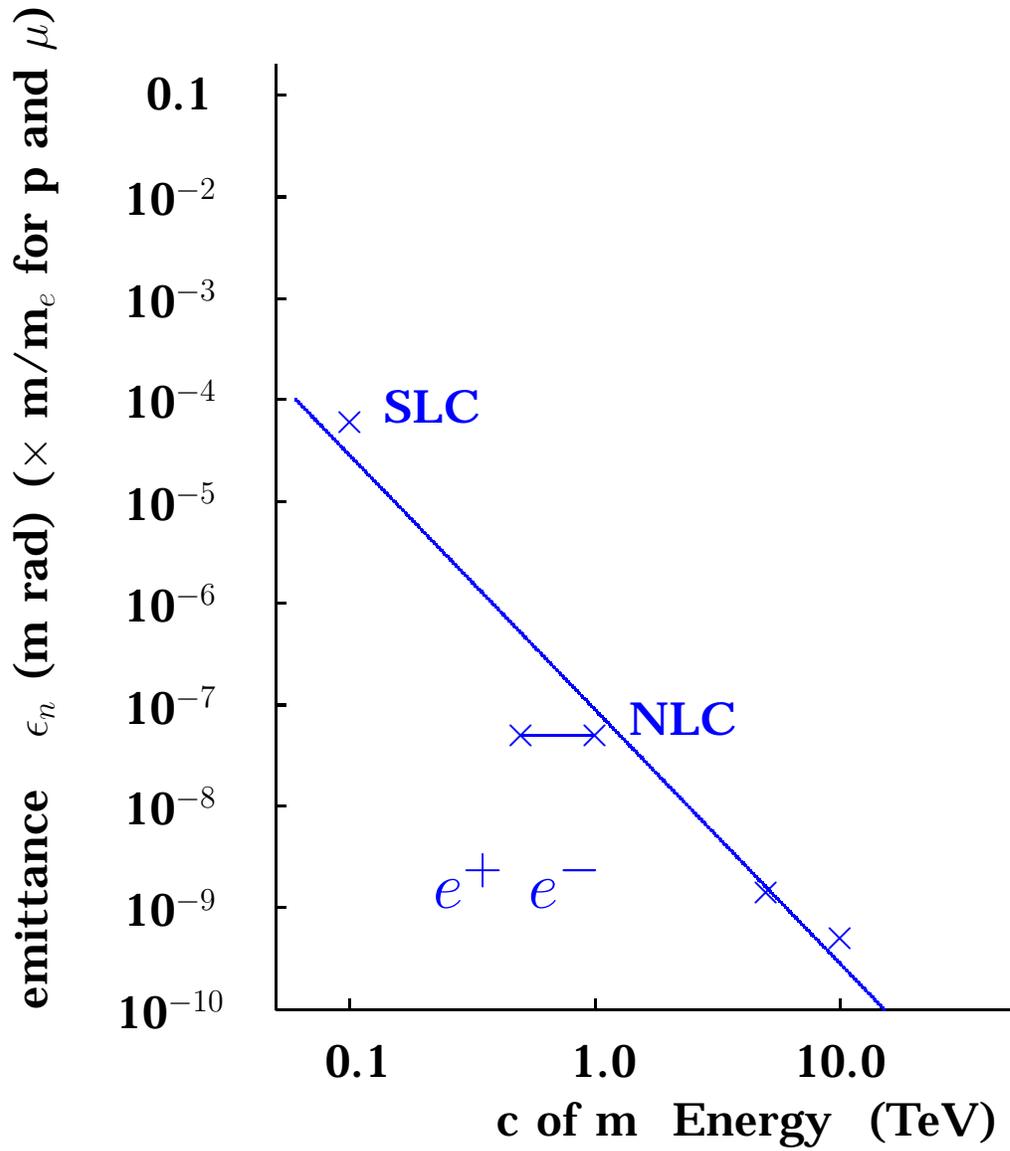
$N/\sigma_x$  is limited by radiation considerations, so

$$\frac{P_{ave}}{\epsilon_n^{(5/7)}} \propto \gamma^3$$

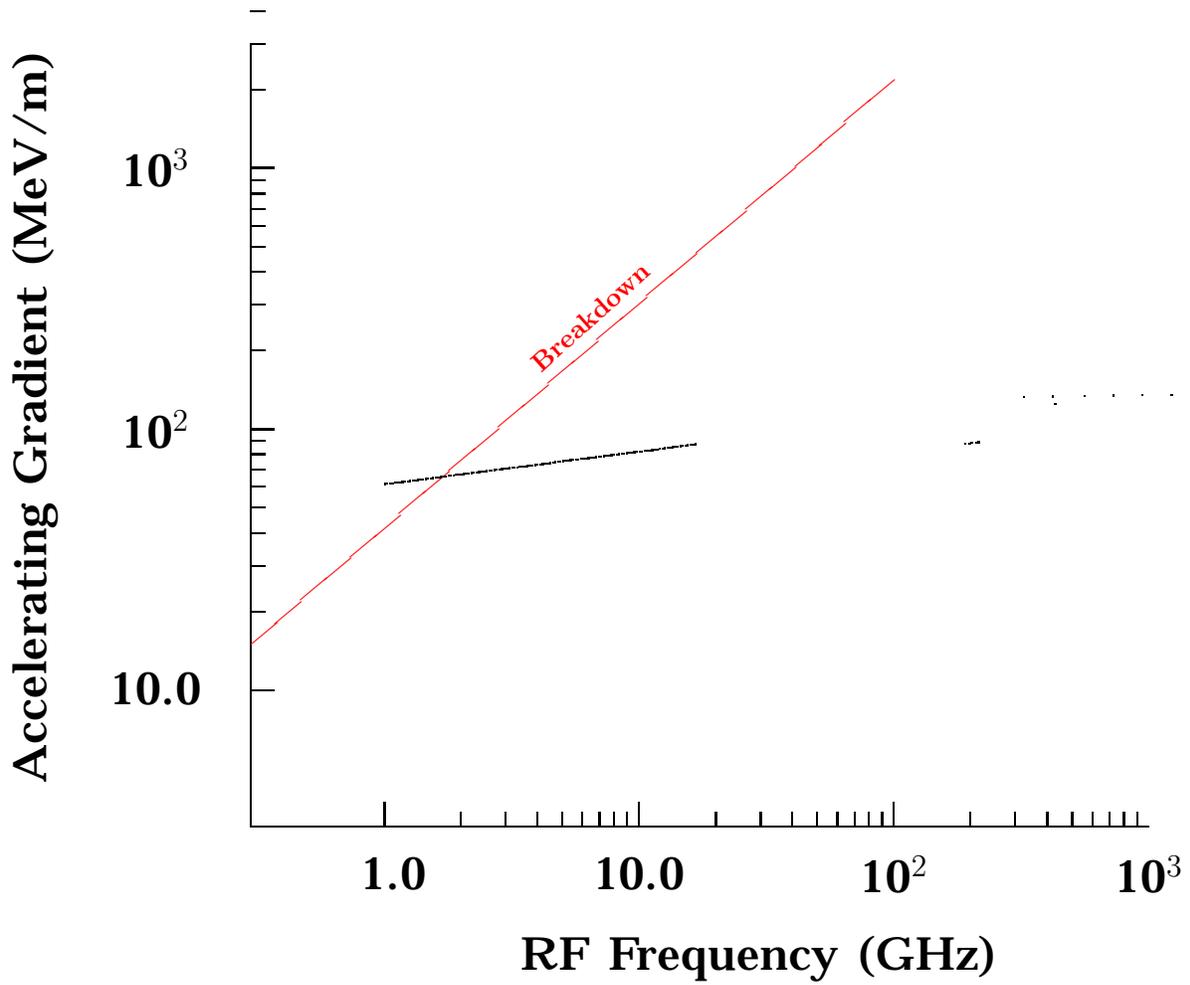
**So we Need:**

- High Beam Power
- Low normalized emittance

# Required Emittances

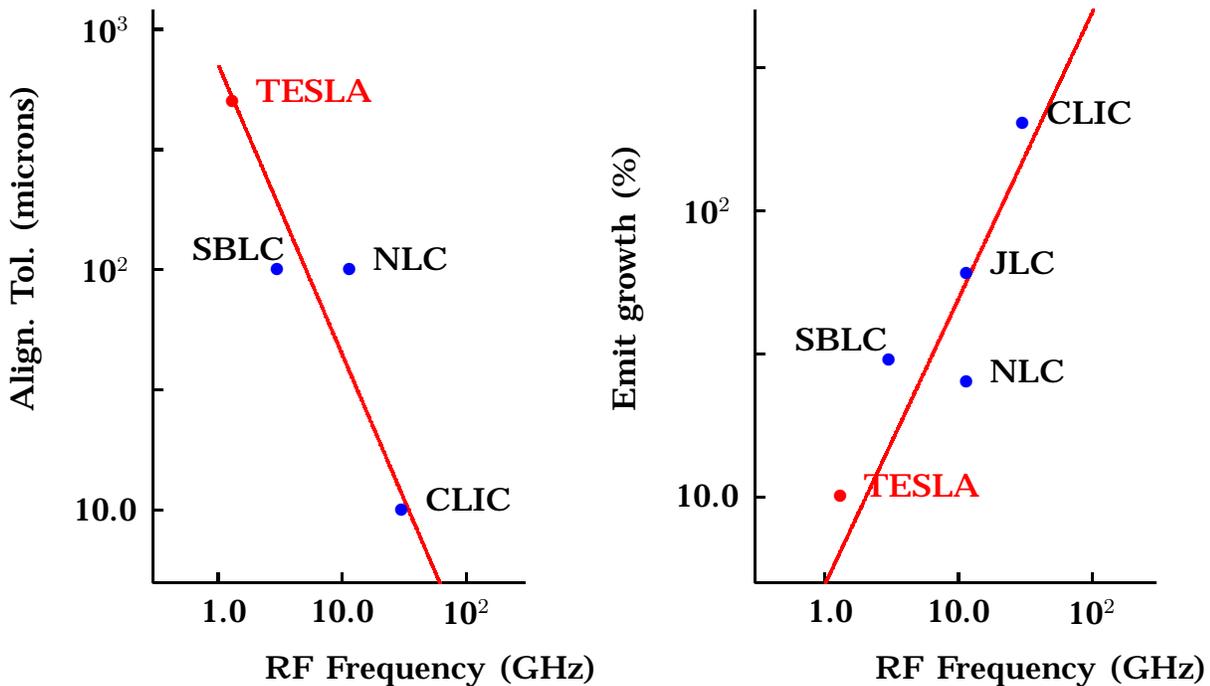


# SIZE AND COST



**But**

**Alignment and emittance growth:**



**High Freq not favored for Lum.**

**Laser and Plasma Acceleration  
will have huge problems with lu-  
minosity**

## Note: with SC cavities (TESLA)

- Easier tolerances
- Low emittance growth
- $2 \times$  wall to beam efficiency
- 1/2 Wall Power
- Better Luminosity at
- Gradient only  $\approx 40\%$  lower

# HADRON COLLIDERS

## Can Costs/TeV be reduced

”Pipeatron” or transmission line magnet (Foster)

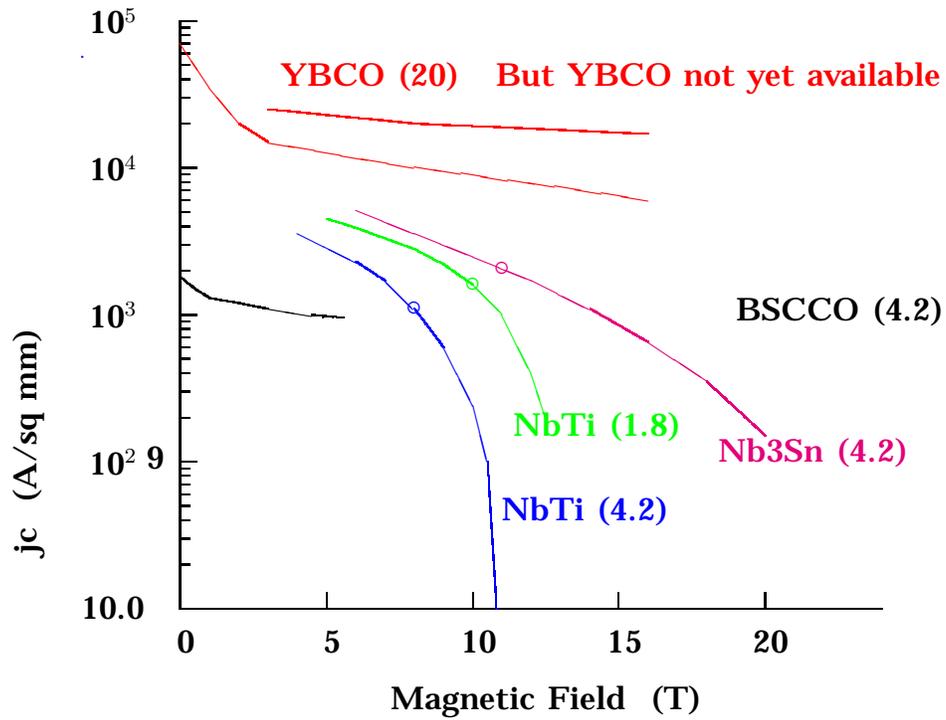
- Linear costs can be reduced
  - improved tunneling methods?
  - low stored magnetic energy
  - low forces
  - very long combined function magnets
- Impedance constrains ( $\propto r^3/B$ )  $\approx 1/30$ 
  - More sophisticated Feedbacks
  - Smooth tube

### Snowmass study:

Cost/TeV of 2T Pipeatron  $\approx$  SSC

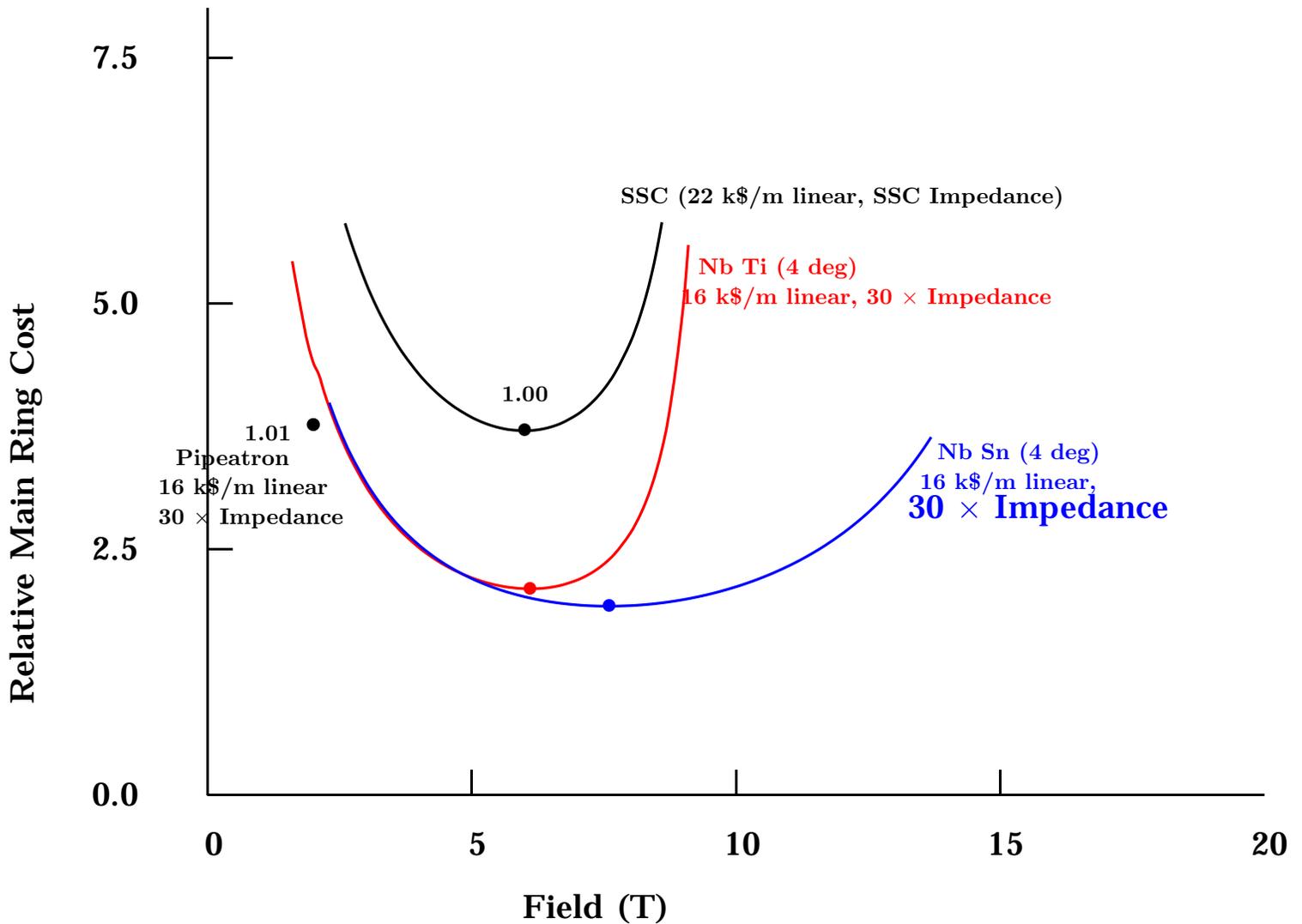
Disappointing

# Will High Field reduce costs ?



# Relative Costs vs. Fields

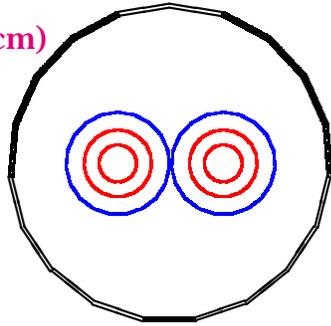
WARNING: Qualitative Only



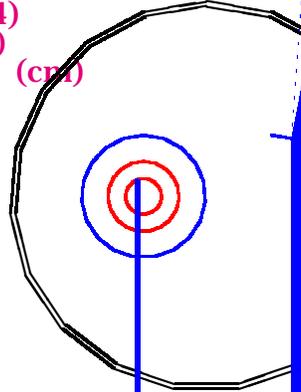
# At Cost Minimums

With SSC costs & LHC Impedance

Nb Ti (1.8)  
B 7.5 (T)  
ap 5.55 (cm)



Nb Sn (4)  
B 9 (T)  
ap 5.27 (cm)



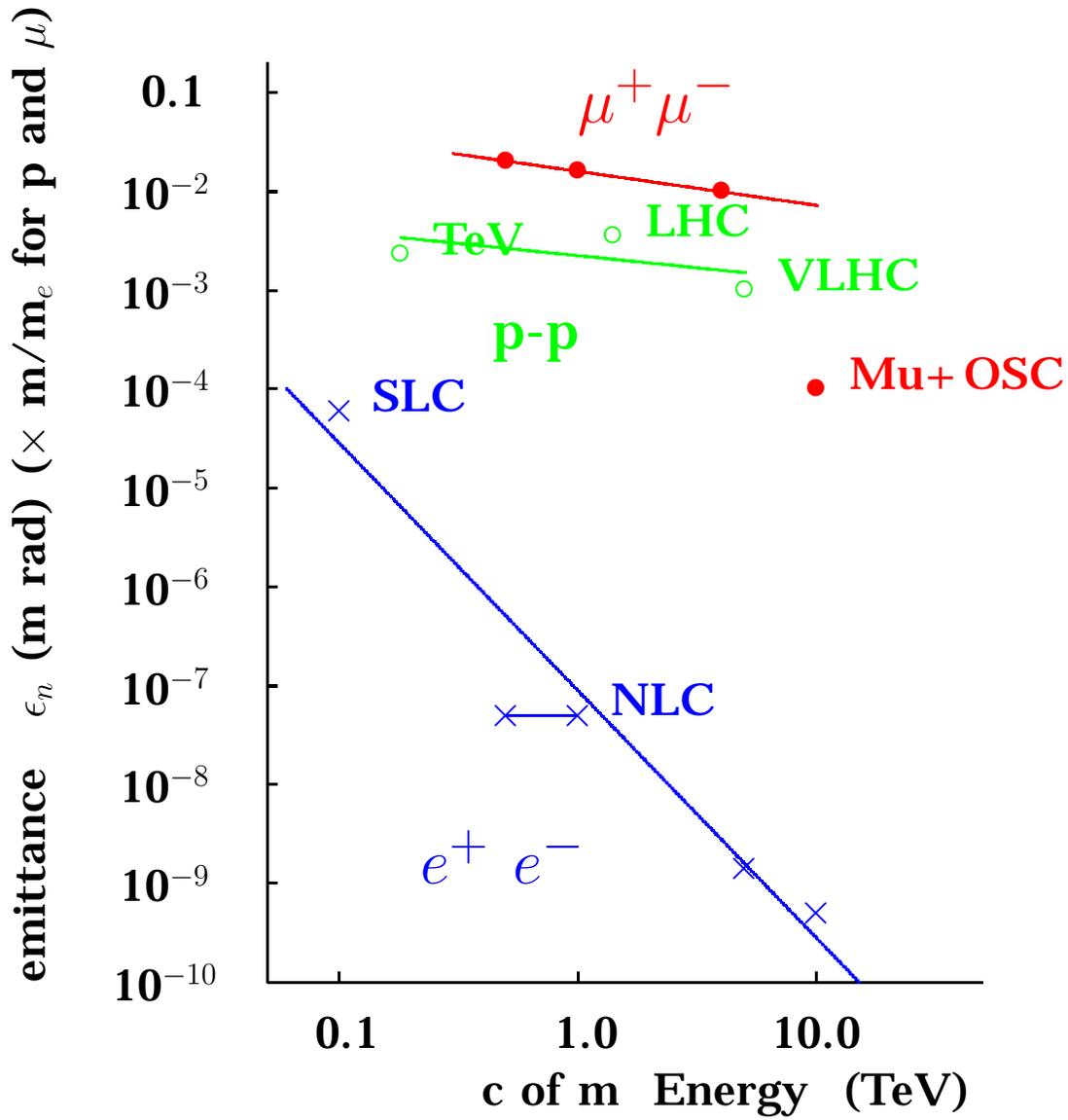
# MUON COLLIDERS

- BNL (USA)
- Cornell (USA)
- Fermi (USA)
- LBNL (USA)
  
- CERN (Switzerland)
- RAL (UK)
  
- KEK (Japan)
  
- BINP (Russia)
  
- + many others

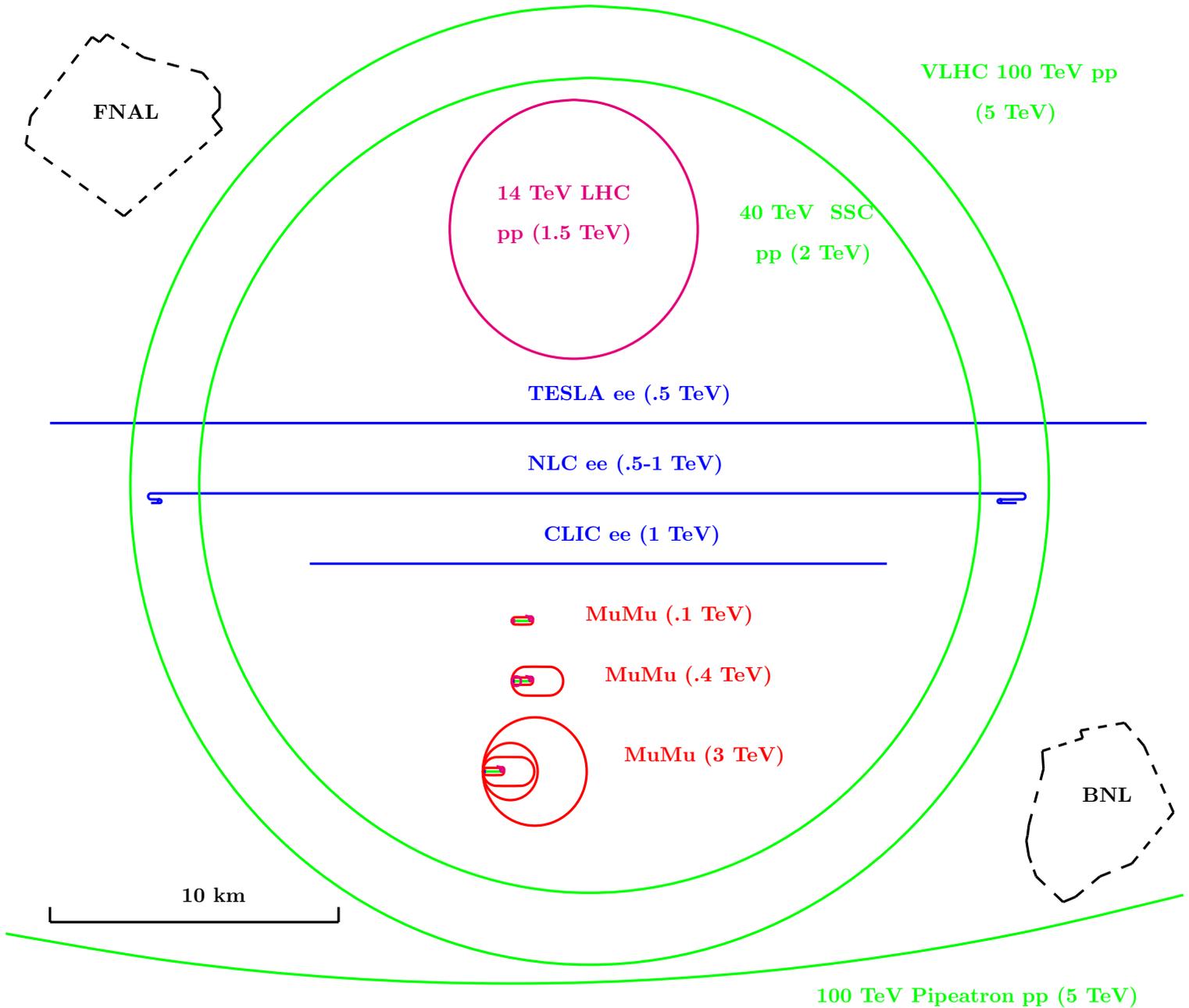
## WHY

- Energy advantage over p's (as for e's)
- Suppressed Synchrotron Radiation ( $\propto \gamma^4 \propto m^{-4}$ )
  - Circular
    - \* smaller
    - \*  $\approx 1000$  turns
  - no "beamstrahlung"
    - \*  $dE/E \rightarrow 0.003\%$
    - \* allows larger N
  - larger spot size & emittance
    - \* easier tolerances
- Direct Higgs Production
  - $\mu^+ \mu^- \rightarrow h, A, H$  ( $\sigma \propto m^2$ )
- $\nu_e$  &  $\nu_\mu$  Neutrino Beams
  - as bi-product, and/or
  - as phase 1

# Easier Tolerances



# Muon Colliders are smaller



## **But: $\mu$ 's in diffuse phase space and $\mu$ 's Decay**

- **Efficient  $\mu$  Production**

- 20 T Solenoid to capture  $\pi$ 's
- Phase Rotation to reduce  $dE/E$

- **Rapid Cooling**

- ionization cooling certain
- study optical and other exotics

- **Rapid Acceleration Required**

- Recirculating Acceleration

- **Beam Pipe Heating**

- 3-6 cm Tungsten Shields

- **Detector Background**

- Loss of Forward Cone

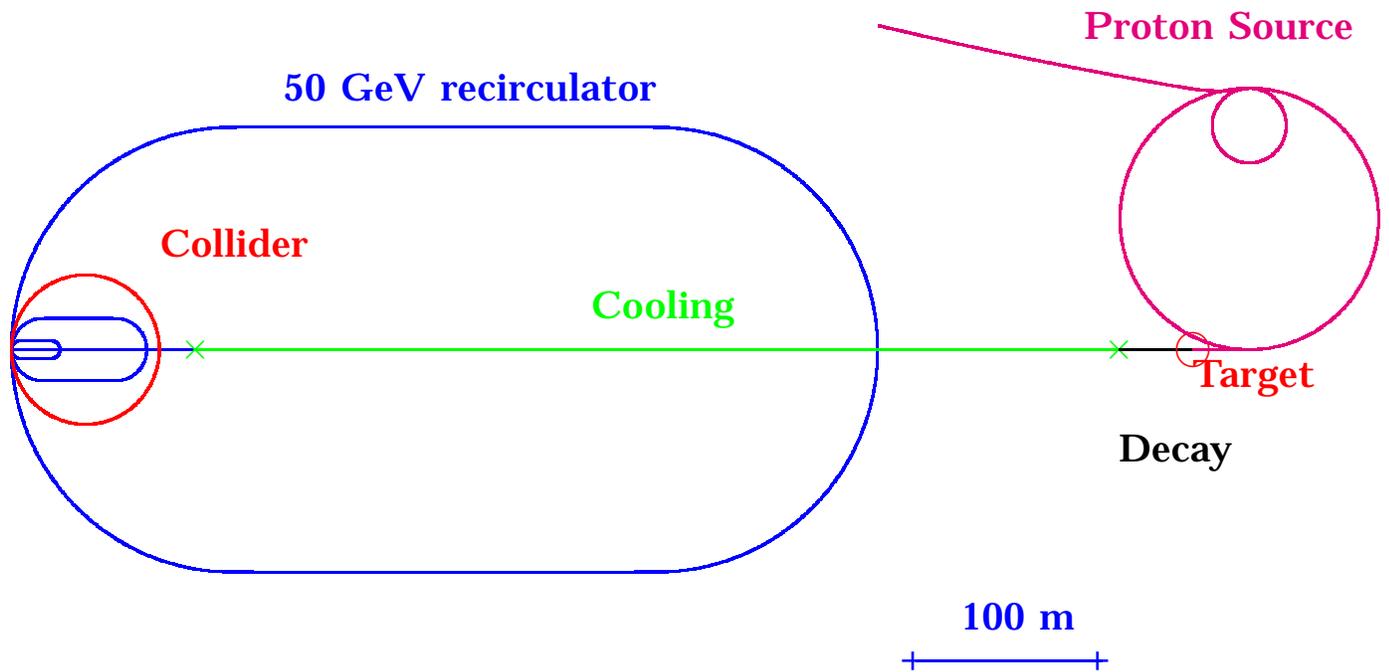
- **Neutrino Radiation**

- Limits Energy to  $\approx 3$  TeV

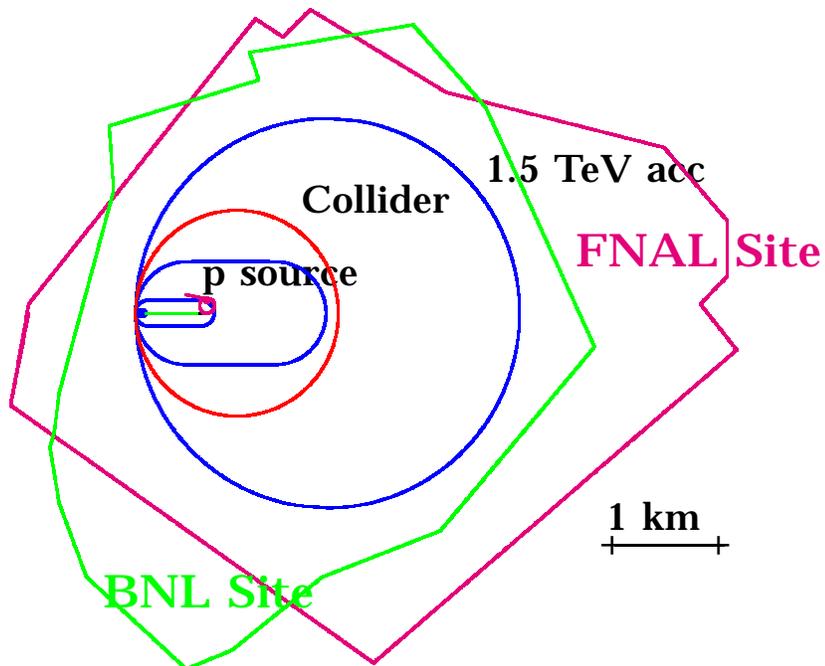
## PARAMETERS

| <b>CoM energy TeV</b>           |                   | <b>3</b>              | <b>0.4</b>            | <b>0.1</b>            |                       |
|---------------------------------|-------------------|-----------------------|-----------------------|-----------------------|-----------------------|
| <i>p</i> energy                 | GeV               | <b>16-24</b>          | <b>16-24</b>          | <b>16-24</b>          |                       |
| <i>p</i> 's/bunch               |                   | $2.5 \times 10^{13}$  | $2.5 \times 10^{13}$  | $5 \times 10^{13}$    |                       |
| Bunches/fill                    |                   | <b>4</b>              | <b>4</b>              | <b>2</b>              |                       |
| Rep. rate                       | Hz                | <b>15</b>             | <b>15</b>             | <b>15</b>             |                       |
| <i>p</i> power                  | MW                | <b>4</b>              | <b>4</b>              | <b>4</b>              |                       |
| $\mu$ /bunch                    |                   | $2 \times 10^{12}$    | $2 \times 10^{12}$    | $4 \times 10^{12}$    |                       |
| $\mu$ power                     | MW                | <b>28</b>             | <b>4</b>              | <b>1</b>              |                       |
| Wall power                      | MW                | <b>204</b>            | <b>120</b>            | <b>81</b>             |                       |
| Collider circ.                  | m                 | <b>6000</b>           | <b>1000</b>           | <b>350</b>            |                       |
| Ave bend B                      | T                 | <b>5.2</b>            | <b>4.7</b>            | <b>3</b>              |                       |
| Depth                           | m                 | <b>500</b>            | <b>100</b>            | <b>10</b>             |                       |
| <b>Rms</b> $\frac{\Delta p}{p}$ | <b>%</b>          | <b>0.16</b>           | <b>0.12</b>           | <b>0.01</b>           | <b>0.003</b>          |
| <b>6D</b> $\epsilon_6$          | $(\pi\text{m})^3$ | $1.7 \times 10^{-10}$ | $1.7 \times 10^{-10}$ | $1.7 \times 10^{-10}$ | $1.7 \times 10^{-10}$ |
| <b>Rms</b> $\epsilon_n$         | $\pi 10^{-6}m$    | <b>50</b>             | <b>50</b>             | <b>85</b>             | <b>290</b>            |
| $\beta^*$                       | cm                | <b>0.3</b>            | <b>2.6</b>            | <b>4.1</b>            | <b>14.1</b>           |
| $\sigma_z$                      | cm                | <b>0.3</b>            | <b>2.6</b>            | <b>4.1</b>            | <b>14.1</b>           |
| $\sigma_r$ spot                 | $\mu m$           | <b>3.2</b>            | <b>26</b>             | <b>86</b>             | <b>294</b>            |
| $\sigma_\theta$ IP              | mrad              | <b>1.1</b>            | <b>1.0</b>            | <b>2.1</b>            | <b>2.1</b>            |
| Tune shift                      |                   | <b>0.044</b>          | <b>0.044</b>          | <b>0.051</b>          | <b>0.015</b>          |
| $n_{turns}$ (e active)          |                   | <b>785</b>            | <b>700</b>            | <b>450</b>            | <b>450</b>            |
| <b>Luminosity</b>               | $cm^{-2}s^{-1}$   | $7 \times 10^{34}$    | $10^{33}$             | $1.2 \times 10^{32}$  | $10^{31}$             |
| Higgs/year                      |                   |                       |                       | $1.9 \times 10^3$     | $3.9 \times 10^3$     |

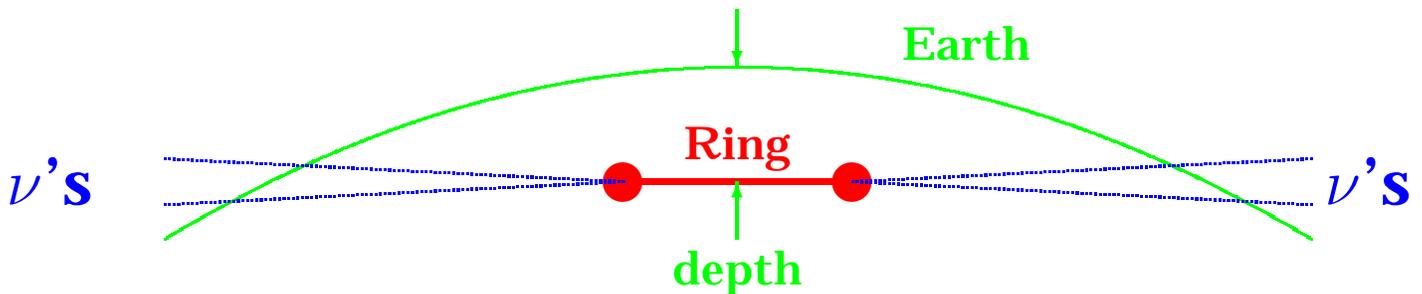
# 100 GeV HIGGS FACTORY



# 3 TeV COLLIDER



# NEUTRINO RADIATION



- Radiation  $\propto E^3/\text{length}^2 \propto E^3/\text{depth}$
- Use: 1/10 Federal limit = 10 mR/year

THEN

- Negligible problem at 1.5 TeV
  - $\approx 1$  mR/year
- E = 3 TeV ok at 300 m depth
  - $\approx 10$  mR/year
- E > 3 TeV Requires:
  - Beam wobbles, and/or
  - Special Locations, and/or
  - Better Cooling (Optical Stochastic?)

## STATUS

- Good theoretical progress

But

- no cooling yet to req. emittance
- Technical challenging

## CURRENT PLANS

- Continue theoretical studies of Muon Collider
- AGS Experiment on Hg (and other) Targets (BNL)
- Testing of RF and absorber systems (FNAL)
- Proposed International Muon Cooling Experiment (MICE)
- Feasibility Studies of " $\nu$  Factories"

# $\mu$ STORAGE RING $\nu$ FACTORY

## ● Study I

- sponsored by Fermi
- finished March 00
- Uses new 16 GeV driver
- "Entry Level" ( $\approx 3 \cdot 10^{19} \mu/10^7 \text{sec}$ )

## ● Study II

- sponsored by BNL
- finished April 01
- Uses upgraded AGS driver
- "Higher Flux" ( $\approx 1.5 \cdot 10^{20} \mu/10^7 \text{sec}$ )

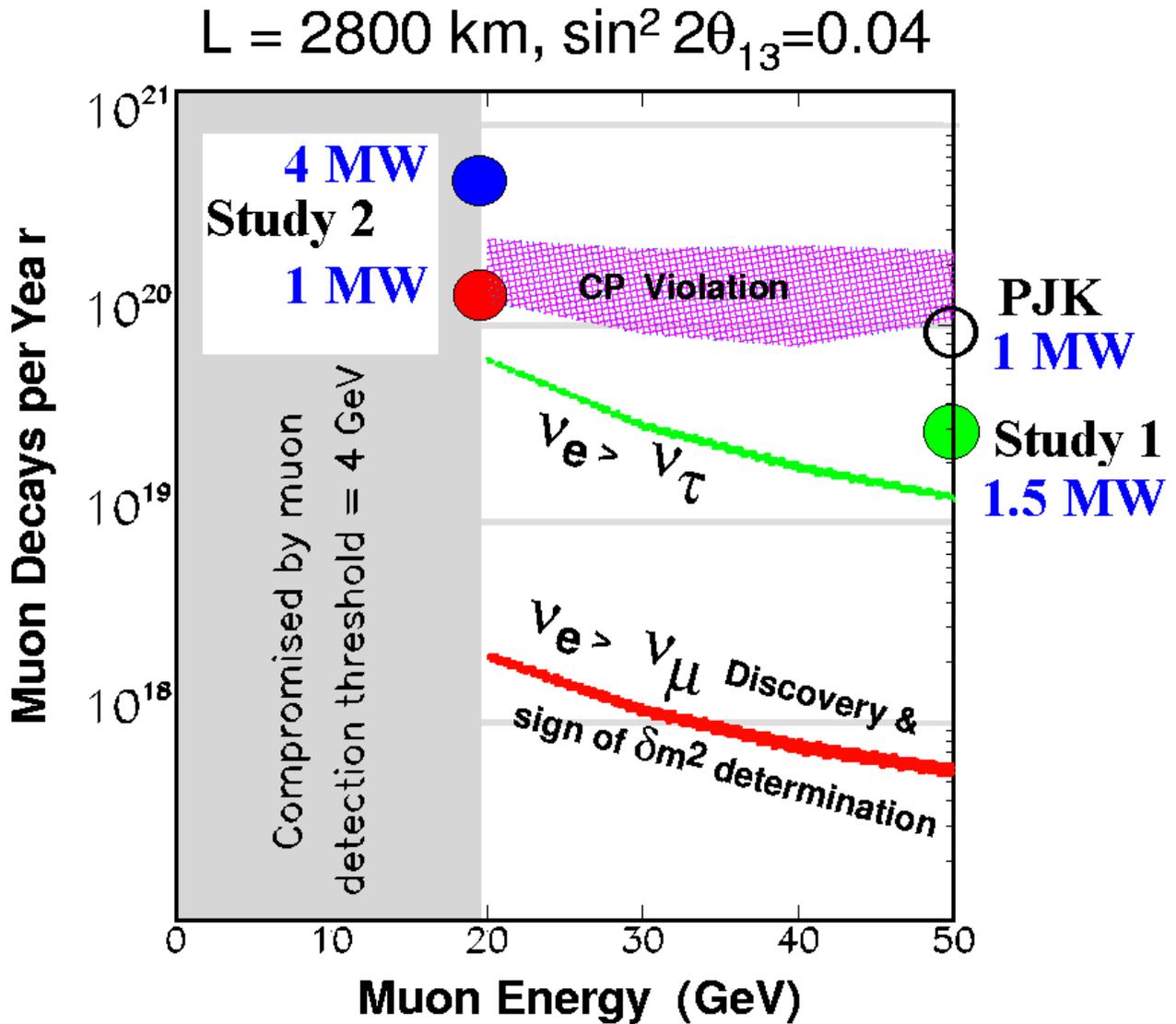
## ● CERN Work

- Uses 2.2 GeV proton linac driver

## ● KEK Work

- Uses 50 GeV JHF driver

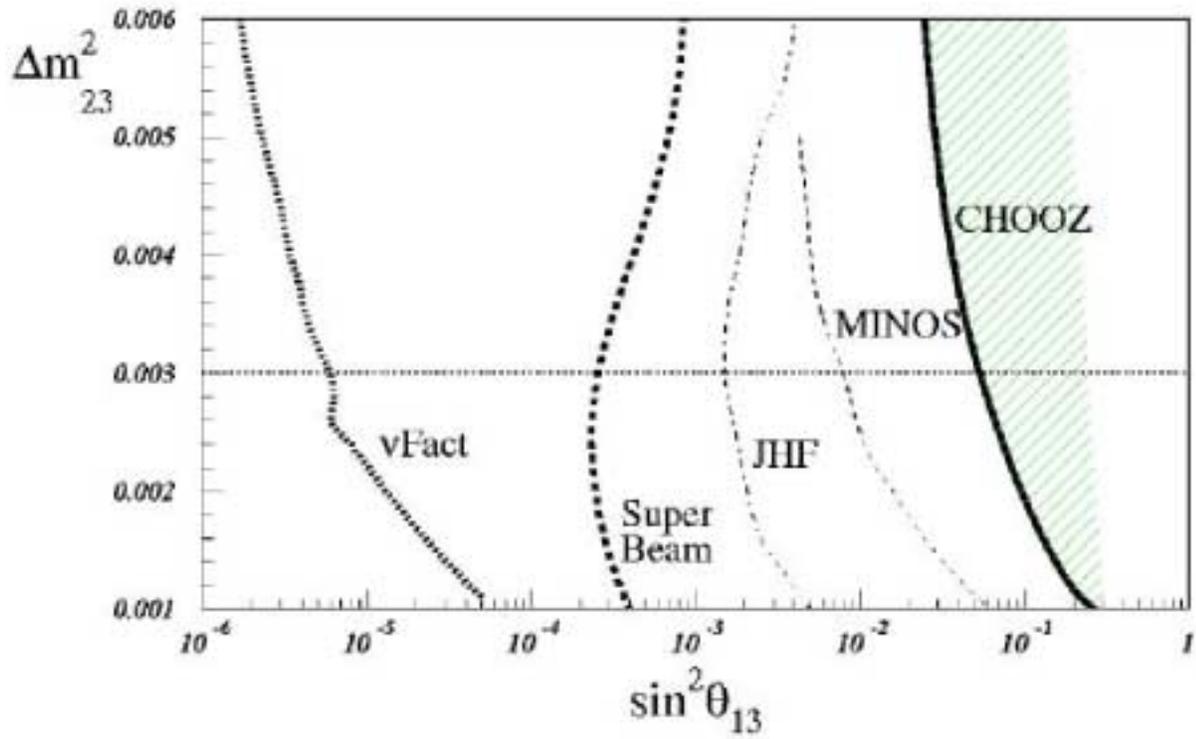
# Physics Reach



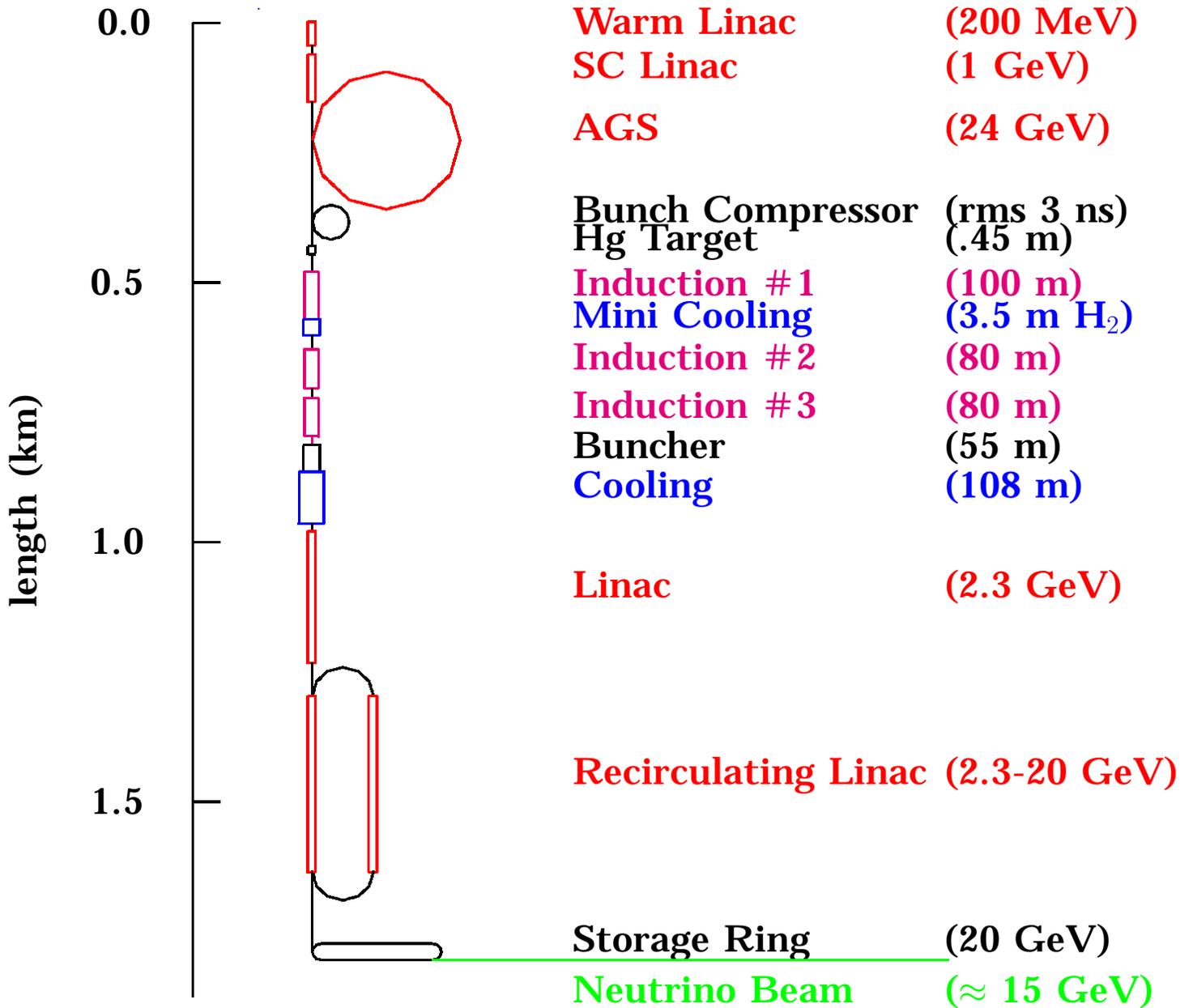
- muon decays in straight section /  $1 \cdot 10^7$  sec
- For Detector mass 50 kT
- Best distance: 2000 - 3000 km

WIPP= 2900 km    Homestake= 2500 km

# Comparisons

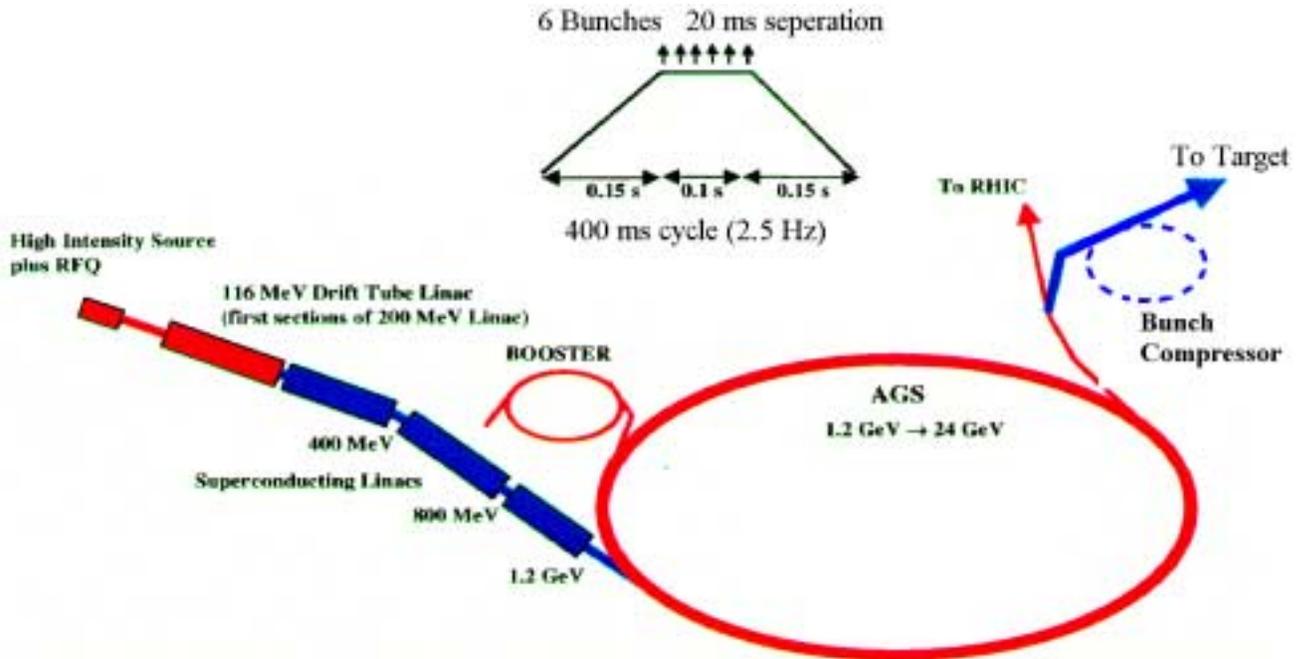


# Schematic



# PROTON DRIVER

- 1 MW BNL AGS Upgrade

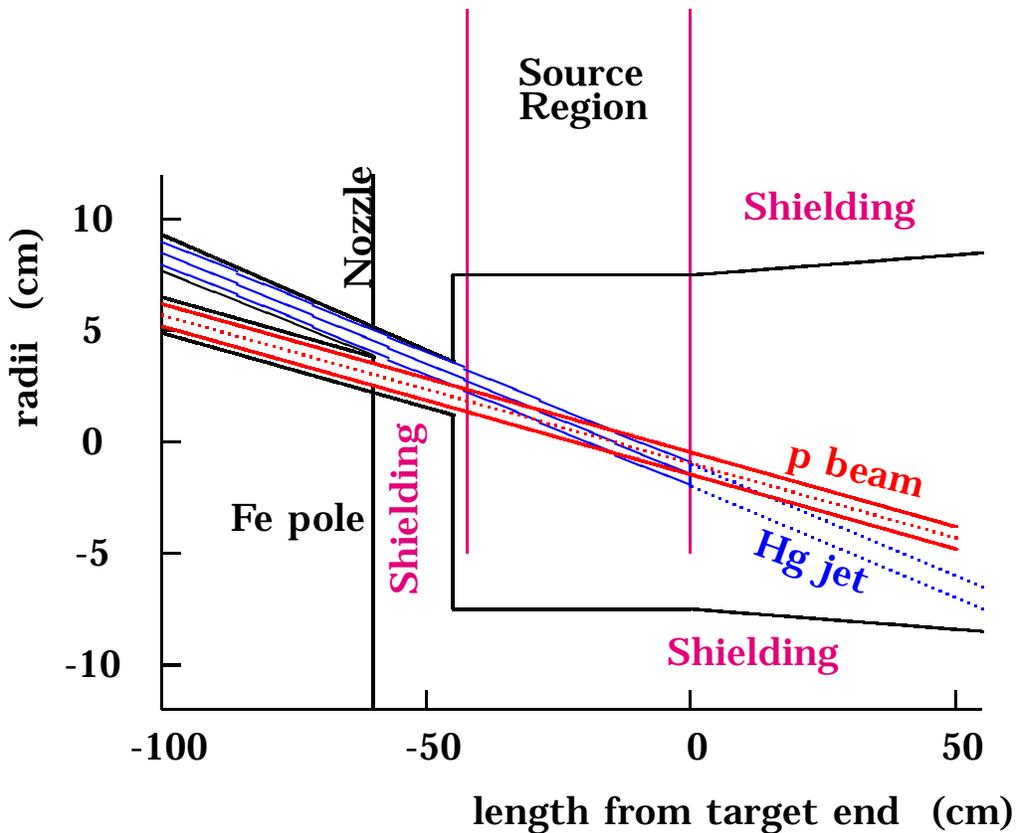


- 4 MW further upgrade
- Similar performance with new 16 GeV Booster at FNAL

# TARGET

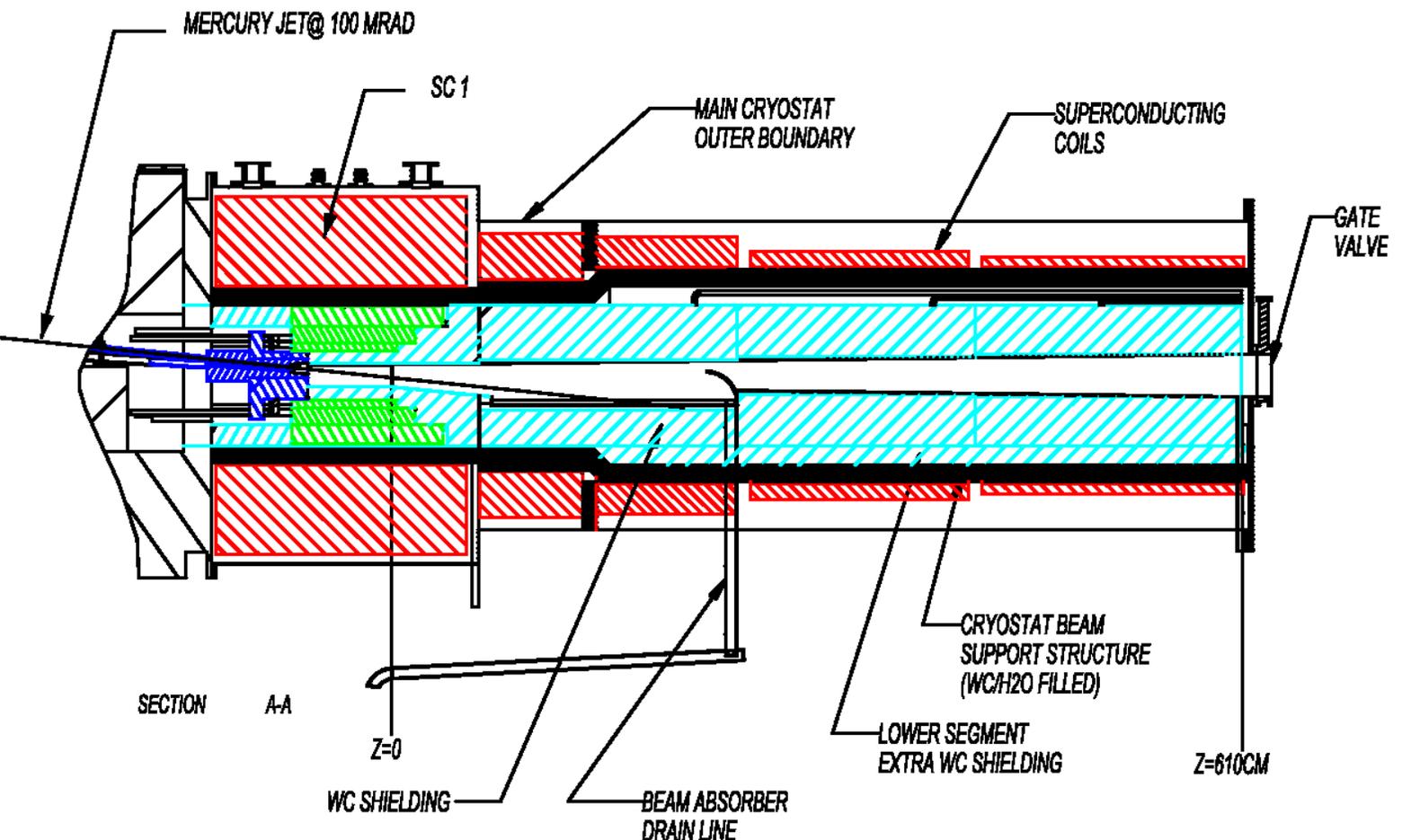
## Mercury jet Target

- $\approx 2 \times$  Carbon (of study 1)
- 20 m/s replaces disturbed
- Nozzle inside field
- OK to 4 MW ?



# Capture Solenoid & Dump

- 20 T hybrid magnet
  - Hollow Conductor Insert
  - Superconducting Outsert
- Taper field to 1.25 T in 18 m
- Mercury pool Beam Dump



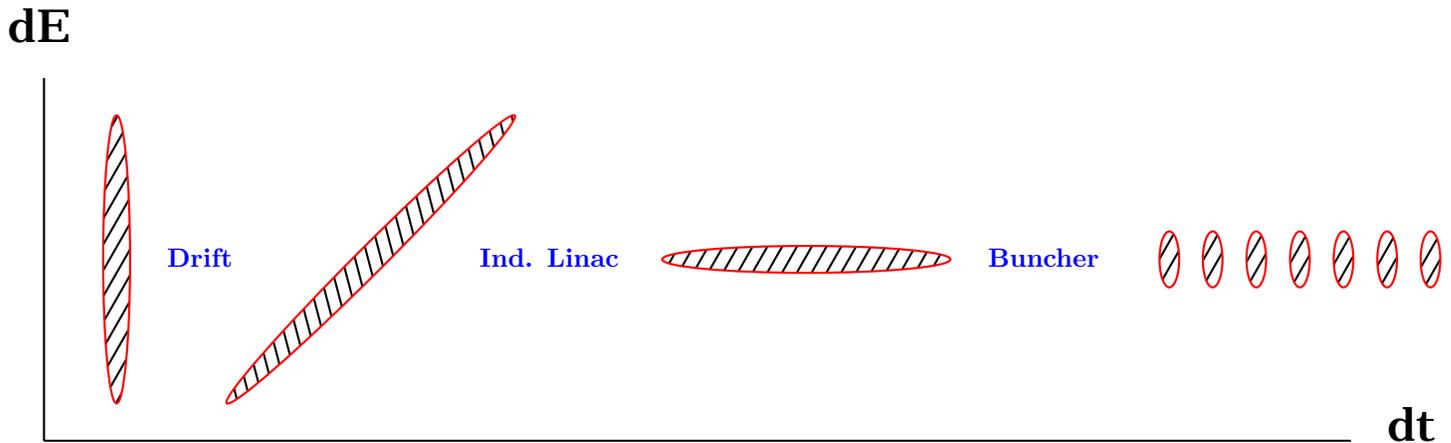
# PROBLEM

- Initial pions have rms  $dp/p \approx 100\%$
- rms Acceptance of cooling  $\approx 8\%$

# SOLUTION:

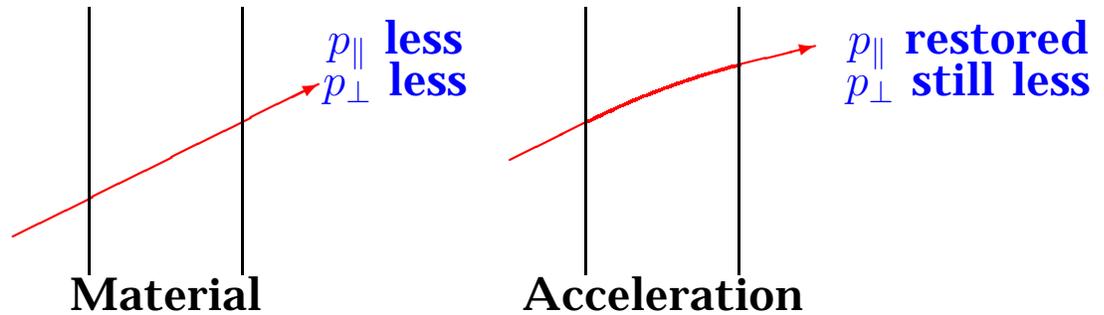
## Phase Rotate & Re-Bunch

- Increase  $dt$
- Decrease  $dE$

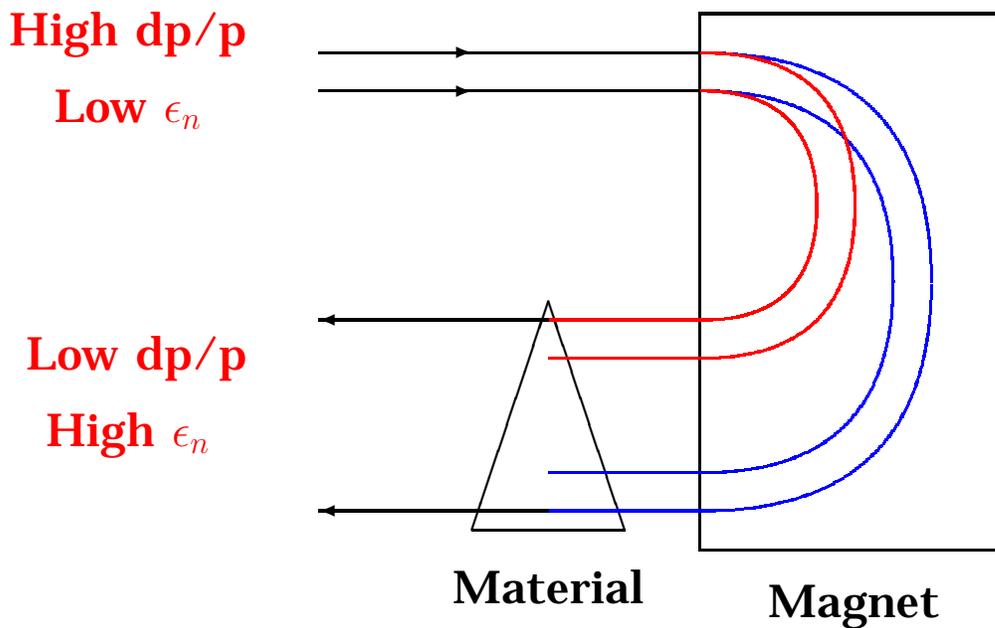


# IONIZATION COOLING

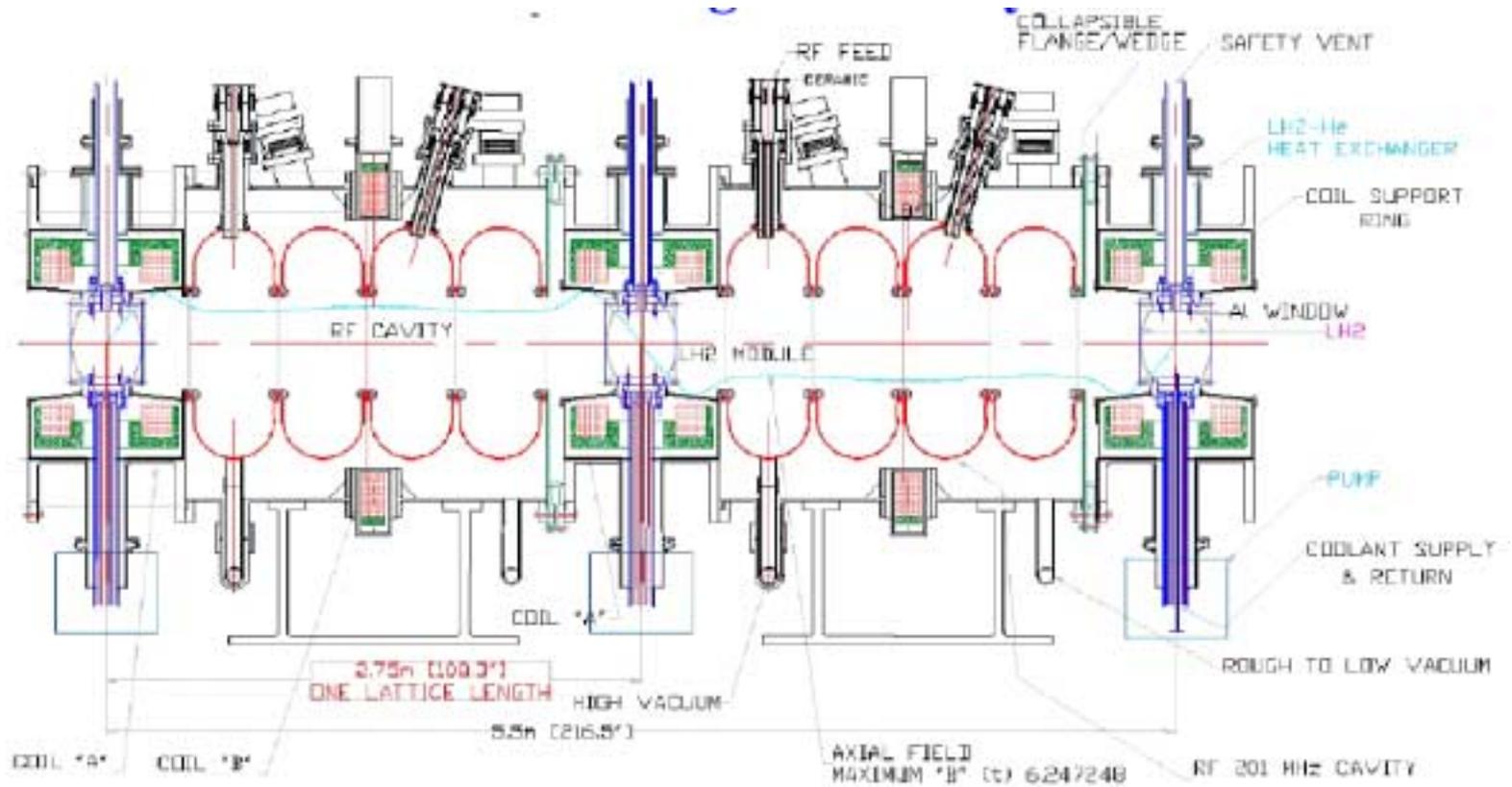
- TRANSVERSE



- $dE/E$  Red by EMIT EXCHANGE



# Cooling Lattice

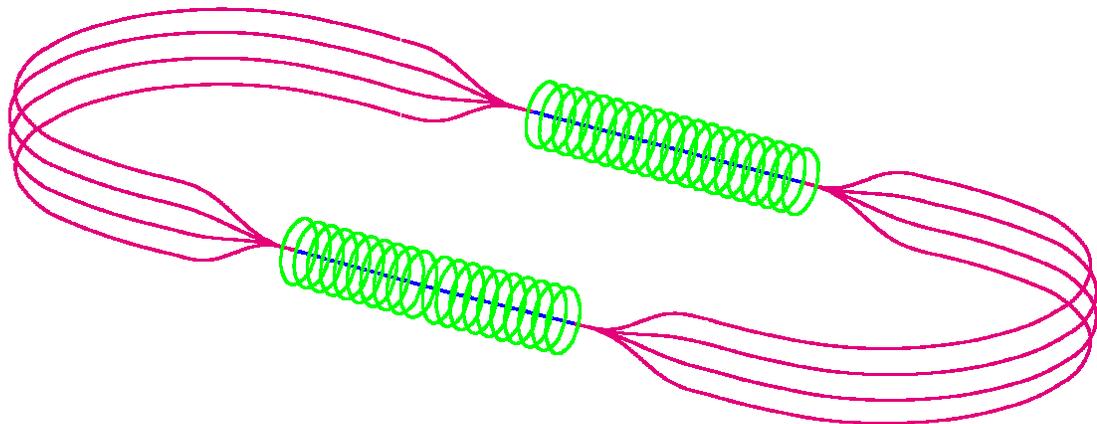


# ACCELERATION

SC Linac followed by  
Recirculating SC Linear Accelerator (RLA)

as at JLAB

in

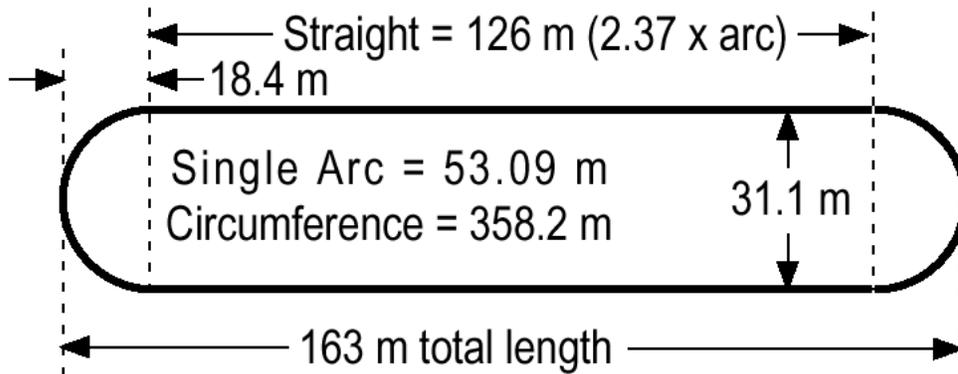


switchyards are passive, using different bending  
for each momentum

# Storage Ring

Arc Magnet Parameters:

|                                |                            |                                   |
|--------------------------------|----------------------------|-----------------------------------|
| $B_1 = 6.93 \text{ T}$ ,       | $G_1 = 0 \text{ T/m}$ ,    | $L_1 = 1.89 \text{ m}$            |
| $B_2 = 0. \text{ T}$ ,         | $G_2 = 35.0 \text{ T/m}$ , | $L_2 = 0.76 \text{ m}$            |
| Average $B = 4.94 \text{ T}$ , |                            | $L_{\text{cell}} = 5.3 \text{ m}$ |



10 Cell Solution  
 $60^\circ$  Arc Cell Phase  
 $\beta_{\text{arc}} = 8.69 \text{ m}$

Empty cell has  
 warm quadrupoles  
 with  $G = 27.2 \text{ T/m}$ .

$$\text{Geometric Decay Ratio} = \frac{126 \text{ m}}{358 \text{ m}} = 0.35 \text{ per straight section}$$

# CONCLUSION

- $e^+ e^-$ 
  - High Grad + High Lum. is hard
  - for  $\leq 800$  GeV SC (TESLA) may be preferred
  - for  $\geq 800$  GeV more R&D required
- $p p$ 
  - No cost reduction with high fields
  - Significant cost reduction with small aperture.
  - R&D Required
- $\mu^+ \mu^-$ 
  - Needs much R&D
- $\mu$  storage ring  $\nu$  Factory
  - "feasible" but expensive
  - Needs R&D to reduce cost